

FYSA2042 osa B, kevät 2019

Koe pe 10.5.2019. Kesto 4 tuntia. Kaavakokoelma lopussa.

Exam Friday May 10th, 2019. Duration: 4 hours. Questions in English and a collection of formulae at the end of the sheet

1. Vastaa lyhyesti seuraaviin kysymyksiin:

- (a) (1p) Miten kemiallinen potentiaali μ liittyy hiukkasluvun odotusarvoon $\langle N \rangle$ suurkanonisessa joukossa, eli miten μ :n nosto tai lasku vaikuttaa odotusarvoon $\langle N \rangle$?
 - (b) (1p) Mitä tarkoitetaan degeneratiolla, kun lasketaan summa yli tilojen? Esimerkiksi mikä on degeneraatio laskettaessa fermioneille summa $\sum_i f(\mathbf{k}_i)$, missä suure f riippuu vain aaltoluvusta, muttei spinistä?
 - (c) (1p) Mitä tarkoittaa Bose-Einstein kondensaatio?
 - (d) (1p) Mitä tarkoitetaan degeneroituneella elektronikaasulla?
 - (e) (2p) Mitä statistisessa mekaniikassa tarkoittaa tasapainotilalla ja toisaalta muuttumatomalla eli stationarisella tilalla?
 - (f) (2p) Mitä kiteen ominaislämmön Debyen mallissa oletetaan fononien spektrista $\epsilon(k)$?
 - (g) (2p) Boltzmannin (kuljetus)teoriassa ratkaistaan jakaumaa $f(\mathbf{r}; \mathbf{p}; t)$. Mitä jakauma kuvaa?
2. (9p) Relativististen elektronien yksihuikkasenergia on $\epsilon_{\mathbf{p}} = cp = \hbar ck$, missä $p = |\mathbf{p}|$. Elektronikaasun lämpötila $T = 0$ ja siinä on N hiukkasta tilavuudessa V . Laske Fermi-energia, sisäenergia ja paine N :n ja V :n funktiona.
 3. (a) (2p) Mitä tarkoittavat keskimääräinen vapaa matka ja keskimääräinen vapaa aika?
(b) (7p) Metalleissa elektronitiheys on usein hyvin korkea, silti elektronien vapaa matka on varsin pitkä. Esim. kuparissa on n. $8,5 \times 10^{28}$ elektronia kuutiometrissä, joten elektronien välinen etäisyys on luokkaa

$$r \approx n^{-1/3} \approx 0,2 \text{ nm} .$$

Silti huoneenlämmössä kuparin elektronien vapaa matka on $\lambda \approx 40 \text{ nm}$. Miksi elektronien vapaa matka metalleissa huoneenlämmössä on paljon suurempi kuin elektronien välinen keskietäisyys? Ja miten lämpötila vaikuttaa asiaan?

4. (10p) Gibbsin entropia on

$$S = -k_B \sum_i p_i \ln p_i ,$$

missä summa on yli mikrotilojen ja p_i on tilan i todennäköisyys. Johda tästä suurkannon jakauman p_i :n lauseke maksimoimalla entropia reunaehdoilla (i) $\sum_i p_i = 1$ (p_i on normitettu todennäköisyysjakauma) (ii) $\sum_i p_i E_i = \langle E \rangle := E$ (termodyn. sisäenergia) ja (iii) $\sum_i p_i N_i = \langle N \rangle := N$ (termod. hiukkasluku).

Vihje: ota reunaehdot Lagrangen kertoimina, maksimoi S ja kirjoita Lagrangen kertoimet lämpötilan ja kemiallisen potentiaalin avulla käyttäen yhtälötä $\sum_i p_i = 1$ ja $dE = TdS - PdV + \mu dN$.

5. (a) (6p) Eristetyn laatikon tilavuus V ja sen sisäseinien lämpötila on T . Osoita, että termodynamiassa tasapainossa laatikossa on keskimäärin

$$N = V \left(\frac{k_B T}{\hbar c} \right)^3 \times a$$

fotonia, missä a on dimensioton luku.

Voit käyttää hyväksi tietoa, että fotonikaasun tilatiheys on $g = V\epsilon^2/(\pi^2\hbar^3c^3)$

- (b) (4p) Osoita edellisen tuloksen avulla, että fotonikaasun ominaislämpö on

$$c_V \propto T^3 .$$

1. Give brief answers to following questions:

- (a) (1p) In the grand canonical ensemble, what is the relation of chemical potential μ and the expectation value of particle number $\langle N \rangle$, that is, how does rising or lowering μ affect the expectation value $\langle N \rangle$?
 - (b) (1p) What is degeneracy in summing over states? For example, what is degeneracy if one calculates the fermion sum $\sum_i f(\mathbf{k}_i)$, where the quantity f depends on the wavenumber, but not on spin?
 - (c) (1p) What is Bose-Einstein condensation?
 - (d) (1p) What is degenerate electron gas?
 - (e) (2p) In statistical mechanics, what does equilibrium state and stationary state mean?
 - (f) (2p) What assumptions about the phonon spectrum $\epsilon(k)$ does the Debye model of lattice specific heat make?
 - (g) (2p) In the Boltzmann (transport) theory one solves the distribution $f(\mathbf{r}; \mathbf{p}; t)$. What does the distribution describe?
2. (9p) Relativistic electrons have single-particle energies $\epsilon_{\mathbf{p}} = cp = \hbar ck$, where $p = |\mathbf{p}|$. The temperature of electron gas is $T = 0$, and there are N electrons in volume V . Calculate the Fermi-energy, internal energy and pressure as functions of N and V .
3. (a) (2p) What does one mean with the mean free path and the mean free time?
- (b) (7p) The electron density in metals is often very high, yet the electron mean free path is quite long. For example, in copper there are about 8.5×10^{28} electrons per cubic meter, so the distance between electrons is about

$$r \approx n^{-1/3} \approx 0.2 \text{ nm} .$$

Still, at room temperature the mean free path of electrons in copper is $\lambda \approx 40 \text{ nm}$. Why is the electron mean free path at room temperature so much longer than the average distance between electrons? What is the role of temperature?

4. Gibbs entropy is

$$S = -k_B \sum_i p_i \ln p_i ,$$

where the sum is over microstates and p_i is the probability of state i . Derive p_i in the grand canonical ensemble by maximizing entropy with boundary conditions (i) $\sum_i p_i = 1$ (p_i is normalized probability distribution) (ii) $\sum_i p_i E_i = \langle E \rangle := E$ (thermod. internal energy), and (iii) $\sum_i p_i N_i = \langle N \rangle := N$ (thermod. number of particles).

Hint: take the boundary conditions as Lagrange multipliers, maximize S , and write the Lagrange multipliers in terms of temperature and chemical potential with the aid of relations $\sum_i p_i = 1$ and $dE = TdS - PdV + \mu dN$.

5. (a) (6p) An isolated box has volume v and the inner wall temperature is T . Show, that in thermodynamical equilibrium there are on average

$$N = V \left(\frac{k_B T}{\hbar c} \right)^3 \times a$$

photons in the box, where a is a dimensionless constant. You may use the fact, that the density of states of photon gas is $g = V\epsilon^2/(\pi^2\hbar^3c^3)$.

- (b) (4p) Using the previous result, show that the specific heat of photon gas is

$$c_V \propto T^3 .$$

Mahdollisesti hyödyllisiä tietoja / potentially useful information

$$\begin{aligned}
& k_B = 1.3805 \times 10^{-23} \text{ J/K} , \quad R = k_B N_A = 8.3143 \text{ J/(mol K)} , \quad N_A = 6.022 \times 10^{23} \text{ 1/mol} \\
& k_B \times 11600 \text{ K} \approx 1 \text{ eV} , \quad 0^\circ\text{C} = 273.15 \text{ K} , \quad 1 \text{ atm} = 1.013 \times 10^5 \text{ Pa} , \quad g = 9.82 \text{ m/s}^2 , \quad c = 2.998 \times 10^8 \text{ m/s} \\
& \hbar = 1.054 \times 10^{-34} \text{ Js} , \quad 1 \text{ J} = 1.602 \times 10^{-19} \text{ eV} , \quad m_e c^2 = 511 \text{ keV} \quad 1u = 1.66 \times 10^{-27} \text{ kg} \\
& dU = \delta Q + \delta W^{\text{rev}} = TdS - PdV , \quad dU = TdS - PdV + \mu dN \\
& F = U - TS , \quad G = U - TS + PV , \quad H = U + PV , \quad \Omega = U - TS - \mu N = -PV \\
& S = k_B \ln \Omega , \quad F = -k_B T \ln Z , \quad U = -\frac{\partial}{\partial \beta} \ln Z , \quad \ln(n!) \approx n \ln n - n , \quad \binom{N}{n} \equiv \frac{N!}{n!(N-n)!} \\
& C_V = T \left(\frac{\partial S}{\partial T} \right)_{V,N} = \left(\frac{\partial U}{\partial T} \right)_{V,N} , \quad C_P = T \left(\frac{\partial S}{\partial T} \right)_{P,N} = \left(\frac{\partial H}{\partial T} \right)_{P,N} , \quad \kappa_T = -\frac{1}{V} \left(\frac{\partial V}{\partial P} \right)_{T,N} , \quad \kappa_S = -\frac{1}{V} \left(\frac{\partial V}{\partial P} \right)_{S,N} \\
& \left(\frac{\partial x}{\partial y} \right)_z = \left[\left(\frac{\partial y}{\partial x} \right)_z \right]^{-1} , \quad \left(\frac{\partial x}{\partial y} \right)_z = \left(\frac{\partial x}{\partial w} \right)_z \left(\frac{\partial w}{\partial y} \right)_z , \quad \left(\frac{\partial x}{\partial y} \right)_z \left(\frac{\partial y}{\partial z} \right)_x \left(\frac{\partial z}{\partial x} \right)_y = -1 \\
& S = -k_B \sum_i p_i \ln p_i , \quad p_i = \frac{1}{Z} e^{-\beta E_i} , \quad Z = \sum_i e^{-\beta E_i} , \quad \beta \equiv \frac{1}{k_B T} \\
& \Omega(T, V, \mu) = -k_B T \ln \mathcal{Z} , \quad p_i = \frac{1}{\mathcal{Z}} e^{-\beta(E_i - \mu N_i)} , \quad \mathcal{Z} = \sum_i e^{-\beta(E_i - \mu N_i)} \\
& \lambda_T = \left(\frac{2\pi\hbar^2}{mk_B T} \right)^{1/2} , \quad Z_1(T, V) = V \lambda_T^{-3} , \quad Z_N(T, V) = \frac{1}{N!} [Z_1(T, V)]^N \\
& PV = Nk_B T = nRT , \quad U = \frac{3}{2} Nk_B T , \quad \left(\frac{dP}{dT} \right)_{\text{ex}} = \frac{\Delta S}{\Delta V} = \frac{L_{1 \rightarrow 2}}{T \Delta V_{1 \rightarrow 2}} \\
& \langle n_i \rangle_{\text{FD}} = \frac{1}{e^{\beta(\epsilon_i - \mu)} + 1} \stackrel{T \rightarrow 0}{\rightarrow} \theta(\epsilon_F - \epsilon_i) , \quad \frac{\partial \langle n_i \rangle_{\text{FD}}}{\partial \epsilon_i} \stackrel{T \rightarrow 0}{\rightarrow} -\delta(\epsilon_i - \epsilon_F) , \quad \langle n_i \rangle_{\text{BE}} = \frac{1}{e^{\beta(\epsilon_i - \mu)} - 1} , \quad \langle n_i \rangle_{\text{MB}} = e^{-\beta(\epsilon_i - \mu)} \\
& \Omega_{\text{FD}} = -k_B T \sum_i \ln[1 + e^{-\beta(\epsilon - \mu)}] \stackrel{T \rightarrow 0}{\rightarrow} -\sum_i (\mu - \epsilon_i) \theta(\mu - \epsilon_i) , \quad \Omega_{\text{BE}} = k_B T \sum_i \ln[1 - e^{-\beta(\epsilon - \mu)}] \\
& \sum_i f(k_i) \approx \frac{gV}{2\pi^2} \int_0^\infty dk k^2 f(k) , \quad \epsilon(k) = \frac{\hbar^2 k^2}{2m} \text{ (NR)} , \quad \epsilon(k) = \hbar ck \text{ (UR)} \\
& u(\omega, T) = \frac{\hbar}{\pi^2 c^3} \frac{\omega^3}{e^{\beta\hbar\omega} - 1} , \quad u(T) = aT^4 , \quad I(T) = \sigma T^4 , \quad a = \frac{\pi^2 k_B^4}{15\hbar^3 c^3} , \quad \sigma = \frac{ac}{4} \\
& \int_0^\infty dx \frac{x^p}{e^x - 1} = p! \zeta(p+1) , \quad \int_0^\infty dx \frac{x^p}{e^x + 1} = (1 - 2^{-p}) \zeta(p+1) \\
& \zeta(s) = \sum_{n=1}^\infty \frac{1}{n^s} , \quad \zeta\left(\frac{3}{2}\right) \approx 2.61 , \quad \zeta(2) = \frac{\pi^2}{6} , \quad \zeta\left(\frac{5}{2}\right) \approx 1.34 , \quad \zeta(3) \approx 1.20 , \quad \zeta\left(\frac{7}{2}\right) \approx 1.13 , \quad \zeta(4) = \frac{\pi^4}{90} \\
& \Gamma(p+1) = p! , \quad \Gamma(z+1) = z\Gamma(z) , \quad \Gamma\left(\frac{1}{2}\right) = \sqrt{\pi} \\
& \int_0^\infty dx e^{-ax^2} = \frac{1}{2} \sqrt{\frac{\pi}{a}} , \quad \int_0^\infty dx x^{2n} e^{-ax^2} = (-1)^n \left[\frac{d^n}{da^n} \frac{1}{2} \sqrt{\frac{\pi}{a}} \right]_{a=1} \\
& \int_0^\infty dx x e^{-ax^2} = \frac{1}{2a} , \quad \int_0^\infty dx x^{2n+1} e^{-ax^2} = (-1)^n \left[\frac{d^n}{da^n} \frac{1}{2a} \right]_{a=1} \\
& \int d\epsilon \frac{f(\epsilon)}{e^{\beta(\epsilon - \mu)} + 1} \approx \int_0^\mu d\epsilon f(\epsilon) + \frac{\pi^2}{6} (k_B T)^2 \left. \frac{df(\epsilon)}{d\epsilon} \right|_\mu + \mathcal{O}(T^4) \\
& \sum_{n=0}^\infty ax^n = \frac{a}{1-x} , |x| < 1 , \quad \sum_{n=1}^\infty \frac{x^n}{n} = -\ln(1-x) , |x| < 1 , \quad e^x = \sum_{n=0}^\infty \frac{x^n}{n!} , \quad (a+b)^n = \sum_{k=0}^n \binom{n}{k} a^k b^{n-k} \\
& \frac{\partial f}{\partial t} + \dot{\mathbf{r}} \cdot \nabla f + \dot{\mathbf{p}} \cdot \nabla_{\mathbf{p}} f = -\frac{1}{\tau} (f - f_0) , \quad \dot{\mathbf{r}} = \mathbf{v} , \quad \dot{\mathbf{p}} = \mathbf{F} \\
& \text{If } \frac{\partial f}{\partial t} = 0 \text{ and } f = f_0 + f' \text{ and } f' \ll f_0, \text{ then } f' \approx -\tau(\dot{\mathbf{r}} \cdot \nabla f_0 + \dot{\mathbf{p}} \cdot \nabla_{\mathbf{p}} f_0)
\end{aligned}$$