



Can we show that cosmological structures are of quantum-mechanical origin?



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Workshop: Inflation and the dark sector — Current challenges and future perspectives

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Inflation

Quantum vacuum fluctuations of the gravitational and scalar fields amplified by gravitational instability and stretched by cosmic expansion



- Strong statement (extraordinary statement requires extraordinary evidence)
- The consequences that can be inferred from this idea are consistent with observations
- This gives an indirect confirmation that cosmological structures have an quantummechanical origin

Any direct evidence?

Inflationary perturbations

$$g_{\mu\nu} = \bar{g}_{\mu\nu}(t) + \hat{\delta g}_{\mu\nu}(t, \boldsymbol{x})$$
$$\phi = \bar{\phi}(t) + \hat{\delta \phi}(t, \boldsymbol{x})$$

Scalar perturbations are described by a single combination of metric and field

 $\widehat{\zeta}(t, \boldsymbol{x})$

fluctuations that directly determines CMB temperature anisotropies

Expansion of Einstein-Hilbert + scalar field action at second order:

$$\hat{H} = \int d^{3}\boldsymbol{k} \begin{bmatrix} \frac{k}{2} \left(\hat{c}_{\boldsymbol{k}} \hat{c}_{\boldsymbol{k}}^{\dagger} + \hat{c}_{-\boldsymbol{k}} \hat{c}_{-\boldsymbol{k}}^{\dagger} \right) & -\frac{i}{2} \underbrace{\left(\hat{a}\sqrt{\epsilon_{1}} \right)^{\prime}}{a\sqrt{\epsilon_{1}}} \left(\hat{c}_{\boldsymbol{k}} \hat{c}_{-\boldsymbol{k}} - \hat{c}_{-\boldsymbol{k}}^{\dagger} \hat{c}_{\boldsymbol{k}}^{\dagger} \right) \end{bmatrix}$$
Creation / annihilation of pairs of particles

Pump field: time-dependent coupling constant Depends only on the scale factor and its derivative Vanishes if *a* is constant

Quantum state of cosmological perturbations

Two-mode squeezed sate



Strongest squeezed state produced in nature (r=50)

Large-squeezing limit: goes to an Einstein-Podolski-Rosen state $|\uparrow,\uparrow\rangle+|\downarrow,\downarrow\rangle$

Can we violate Bell's inequalities with the CMB?



highly-non classical state?

Classicality in the Wigner approach

Wigner function
$$W(q,p) = \int \Psi^* \left(q - \frac{u}{2}\right) e^{-ipu} \Psi \left(q + \frac{u}{2}\right) \frac{\mathrm{d}u}{2\pi}$$

Evolution equation: $\frac{\partial}{\partial t}W(q, p, t) = -\{W(q, p, t), H(q, p, t)\}_{\text{Poisson bracket}}$ for quadratic Hamiltonians



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Weyl Transform
$$\widetilde{A}(q,p) = \int du e^{-ipu} \left\langle q + \frac{u}{2} \middle| \hat{A} \middle| q - \frac{u}{2} \right\rangle$$
(with this definition: $W = \frac{\widetilde{\rho}}{2\pi}$)Expectation value of quantum operators $\langle \hat{A} \rangle = \int \widetilde{A}(q,p) W(q,p) dq dp$

W > 0 — "quasi-probability distribution"

John Bell 1986, *EPR correlations and EPW distributions*: Bell inequality violation requires non-positive Wigner function

Henderson and Vedral 2001; Ollivier and Zurek 2001



Idea: Find two ways to calculate the mutual information between A and B that coincide for classical correlations but may differ in quantum systems

$$\mathcal{I} = S(A) + S(B) - S(A, B)$$

 $\mathcal{J} = S(A) - S(A|B)$ with respect to measurements $\hat{\Pi}_j$

$$\hat{\Pi}_{j} : \text{complete set of projectors defined on } \mathcal{E}_{B} \hat{\rho} \to \hat{\rho} \hat{\Pi}_{j} / p_{j} \text{ with probability } p_{j} = \text{Tr} \left(\hat{\rho} \hat{\Pi}_{j} \right) \text{ and } \rho_{A;\hat{\Pi}_{j}} = \text{Tr}_{B} \left(\rho \hat{\Pi}_{j} / p_{j} \right) S(A|B) = \sum_{j} p_{j} S \left(\rho_{A;\hat{\Pi}_{j}} \right)$$

$$\delta(A,B) = \min_{\{\hat{\Pi}_j\}} \left(\mathcal{I} - \mathcal{J} \right)$$

Example:
$$|\Psi\rangle = \frac{|\downarrow\downarrow\rangle + |\uparrow\uparrow\rangle}{\sqrt{2}}$$

 $\rho = \frac{1}{2}|\downarrow\downarrow\rangle\langle\downarrow\downarrow| + \frac{1}{2}|\uparrow\uparrow\rangle\langle\uparrow\uparrow| + \frac{1}{2}|\downarrow\downarrow\rangle\langle\uparrow\uparrow| + \frac{1}{2}|\uparrow\uparrow\rangle\langle\downarrow\downarrow|$

$$\begin{array}{l} \text{Example:} \ |\Psi\rangle = \frac{|\downarrow\downarrow\rangle + |\uparrow\uparrow\rangle}{\sqrt{2}} \\ & \downarrow \\ & \downarrow \\ & \rho = \frac{1}{2}|\downarrow\downarrow\rangle\langle\downarrow\downarrow| + \frac{1}{2}|\uparrow\uparrow\rangle\langle\uparrow\uparrow| + \frac{z}{2}|\downarrow\downarrow\rangle\langle\uparrow\uparrow| + \frac{z}{2}|\uparrow\uparrow\rangle\langle\downarrow\downarrow| \end{array}$$





For the two-mode squeezed state of inflation: J. Martin, V.V., 1510.04038

$$\delta(\boldsymbol{k}, -\boldsymbol{k}) = \cosh^2 r_k \log_2 \left(\cosh^2 r_k \right) \\ - \sinh^2 r_k \log_2 \left(\sinh^2 r_k \right)$$

 $\sim 150\,$ at the end of inflation

So is the CMB very classical or very quantum?



You have reached the point of maximum confusion



Revzen 2006

Bell inequalities can be violated even when W>0 with improper operators

Proper operator: \widetilde{A} takes values within the spectrum of \widehat{A} .

Bell inequalities



- Bipartite system: *k* and -*k*
- Entangled system: two-mode squeezed state
- improper, spin-like operators

Bell inequalities



continuous variable

$$\hat{q}_{\boldsymbol{k}} = \frac{\hat{c}_{\boldsymbol{k}} + \hat{c}_{\boldsymbol{k}}^{\dagger}}{\sqrt{2k}} = \hat{q}_{\boldsymbol{k}}^{\dagger}$$



- Divide the real axis into intervals $[n\ell, (n+1)\ell]$
- Perform a measurement of q_{k}
- Return $S_z(\ell) = (-1)^n$

Larsson 2004

04
$$\hat{S}_{z}(\ell) = \sum_{n=-\infty}^{\infty} (-1)^{n} \int_{n\ell}^{(n+1)\ell} \mathrm{d}q_{k} |q_{k}\rangle \langle q_{k}| \longrightarrow \hat{S}_{z}^{2}(\ell) = 1$$

$$\hat{S}_{+}(\ell) = \sum_{n=-\infty}^{\infty} (-1)^{n} \int_{2n\ell}^{(2n+1)\ell} \mathrm{d}q_{k} |q_{k}\rangle \langle q_{k} + \ell| \underbrace{\qquad} \hat{S}_{x}(\ell) = \hat{S}_{+}(\ell) + \hat{S}_{+}^{\dagger}(\ell) \\ \hat{S}_{y}(\ell) = -i \left[\hat{S}_{+}(\ell) - \hat{S}_{+}^{\dagger}(\ell) \right]$$

 $\longrightarrow \left[\hat{S}_{i}\left(\ell\right),\hat{S}_{j}\left(\ell\right)\right]=2i\epsilon_{ijk}\hat{S}_{k}\left(\ell\right) \quad \text{obey spin algebra}$

Bell inequalities in the CMB



Bell inequalities in the CMB



Can we detect quantum correlations using "position" measurements only?

 $\widetilde{\zeta}_{k} = \zeta_{k}$ and $\widetilde{f}(\zeta_{k}) = f(\zeta_{k})$ so according to Revzen's theorem: not with Bell inequalities!

Leggett-Garg inequalities

Two-time correlators $C_{ab} = \left\langle \hat{S}_z(t_a, \ell) \hat{S}_z(t_b, \ell) \right\rangle$

Leggett-Garg three strings $K_3 = C_{ab} + C_{bc} - C_{ac}, \quad K'_3 = -C_{ab} - C_{bc} - C_{ac}$

Classically: $-3 \le K_3, K_3' \le 1$



But requires to measure zeta at three different times ...

Non-discordant states

Can we detect quantum correlations using single-time, "position" measurements only? Classical states = non-discordant states $\delta(\mathbf{k}, -\mathbf{k}) = 0$



Non-discordant state sharing the same two-point functions as the two-mode squeezed state

Non-discordant states

Can we detect quantum correlations using single-time, "position" measurements only?

Classical states = non-discordant states $\delta(\mathbf{k}, -\mathbf{k}) = 0$

Theorem: the only classical Gaussian states are product states

Adesso, Datta 2010; Rahimi-Keshari, Calves, Ralph 2013; Mista, Mc Nulty 2014



Conclusions

- Cosmological perturbations are placed in a two-mode highly squeezed state in the very early Universe
- Such a state has a large **quantum discord**, denoting the presence of **large quantum correlations** between particles created with opposite wave momenta
- In principle, **Bell experiments** can therefore be constructed that would prove that CMB anisotropies are of quantum mechanical origin
- Legget-Garg inequalities evade this issue but require to measure perturbations at different times
- The CMB cannot have been placed in a **classical Gaussian state**. Current constraints on non-Gaussianities may be already sufficient to exclude non-discordant states!
- Role of decoherence?

Decoherence



Decoherence



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