

# Spectral rigidity of planets

Quasilinear Equations, Inverse Problems and Their Applications

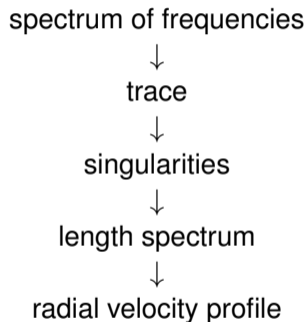
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In collaboration with:  
M. V. de Hoop & V. Katsnelson

24 August 2021

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## 1 The idea

- The spectrum of eigenfrequencies
- The trace
- The length spectrum
- Linearization
- The result

## 2 Mathematics

## 3 Physics

## 4 The Herglotz condition

# The spectrum of eigenfrequencies

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- The amplitudes of the various modes depend on the event, the frequencies do not.
- The set of frequencies is the spectrum of free oscillations.
- These frequencies are (square roots of) eigenvalues of an elliptic partial differential operator, “the elastic Laplacian”.



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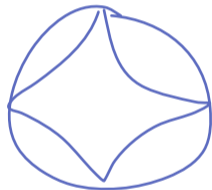
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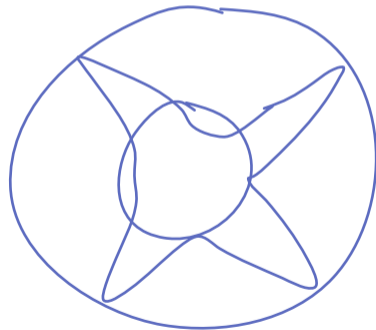


- The function  $t \mapsto \mathrm{tr}(\partial_t G)$  is mostly smooth. It has a singularity at  $t = T$  if and pretty much only if  $T$  is the length of a periodic ray path (periodic broken geodesic).

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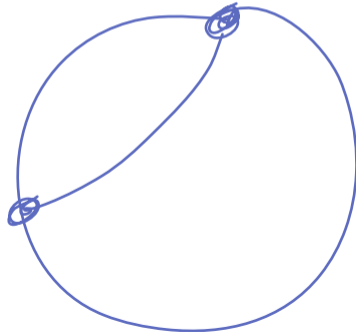
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- At interfaces the rays can choose to reflect or refract (and change polarization).
- The length spectrum is computable from the eigenfrequencies (singularities of the trace) but can also be measured directly. (Both appear possible on Mars, in principle.)
- Using the length spectral data is like travel time tomography but without endpoints.

# Linearization

- The full problem is non-linear and hard, so we linearize.

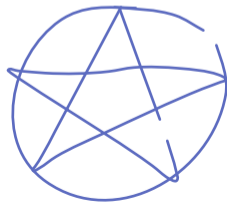
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- Linearizing normal travel time tomography between surface points leads to geodesic X-ray tomography.
  - This problem is well understood on some manifolds.
- Linearizing periodic travel time tomography leads to periodic broken ray tomography.
  - This problem is beginning to be understood on very few manifolds.



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A spherically symmetric planet with radial velocity profile(s) with **jump discontinuities** and mild geometric conditions is determined by its spectrum or length spectrum.



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A spherically symmetric planet with radial velocity profile(s) with jump discontinuities and mild geometric conditions is determined by its spectrum or length spectrum.

≈ Earth-like planets are locally uniquely determined by the spectrum of free oscillations.

# Outline

1 The idea

2 Mathematics

- Spherical symmetry
- Polarization
- Basic rays
- Conjugate points
- The theorems

3 Physics

4 The Herglotz condition

# Spherical symmetry

- For a Riemannian metric (elliptically anisotropic wave speed) on a ball there are two concepts of spherical symmetry:

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- These are equivalent — up to changing coordinates.

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- In spherical symmetry one can calculate explicitly rather than just qualitatively.
- The geodesic flow is integrable with conserved quantities.
- Geometry behaves well under the Herglotz condition:  $\frac{d}{dr} \left( \frac{r}{c(r)} \right) > 0$  for all  $r > 0$ .

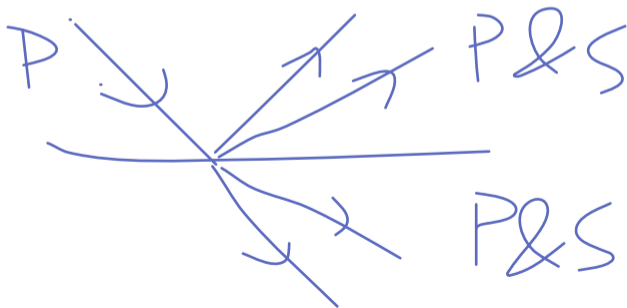
# Polarization

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- The results hold for the isotropic elastic system (2 radial Lamé parameters) but are trickier to state.
- Waves split 4 ways at interfaces.
- In the scalar case the spectrum of free oscillations is modeled geometrically “simply” by the Neumann spectrum of a Laplace–Beltrami operator.

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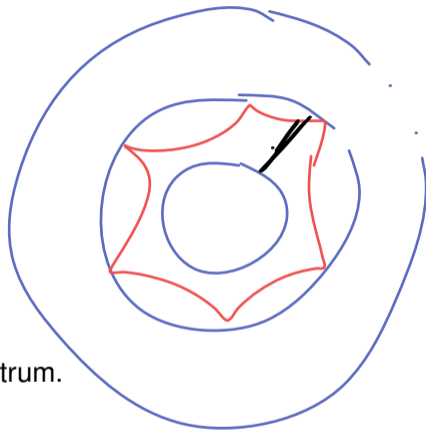
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# Basic rays

We define a periodic broken ray to be *basic* if

- it stays in a single layer,
- it stays in the same polarization, and
- it is of one of two types:
  - It only reflects from the top interface.
  - It is radial.

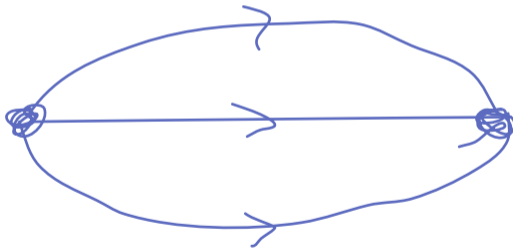
The lengths of these constitute the basic length spectrum.



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- Conjugate points cause trouble in geometric proofs!
- We make two assumptions:
  - Countable conjugacy condition:  
There are only countably many conjugate points (up to rotations and shifts).
  - Periodic conjugacy condition:  
No point is self-conjugate along a basic ray.

# The theorems

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A parameter  $\tau \in (-\varepsilon, \varepsilon)$  describing a family of models.

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# The theorems

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Earlier result without discontinuities: [arXiv:1705.10434](https://arxiv.org/abs/1705.10434)

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- 2 Mathematics
- 3 Physics
  - Planetary models
  - Measurability
- 4 The Herglotz condition

# Planetary models

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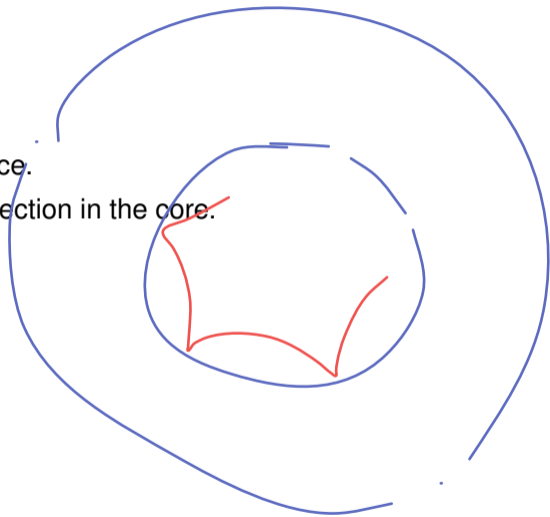
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- Is Herglotz a consequence of physics?



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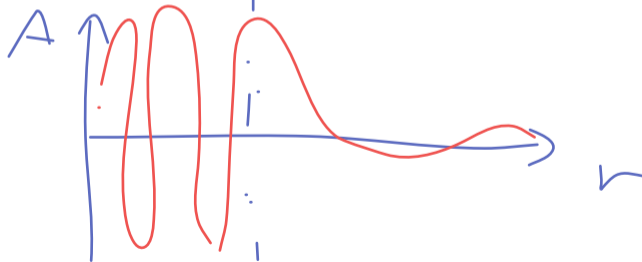
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- In the high frequency limit the amplitudes and the relative surface amplitudes go down.
- Noise is a problem.



- 1 The idea
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- 3 Physics
- 4 The Herglotz condition
  - The classical Herglotz condition
  - Reduced phase space
  - Herglotz as convexity

# The classical Herglotz condition

$$\rho(r) := \frac{r}{c(r)}$$

$$\rho'(r) > 0$$

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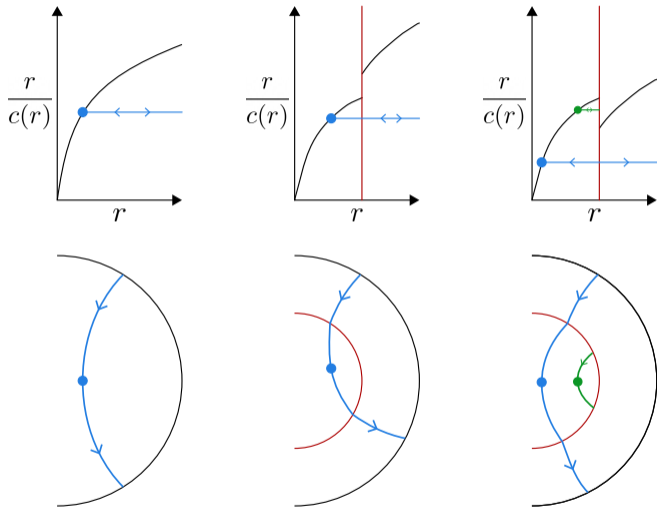
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- Properties:
  - $L$  is conserved — also at interfaces.
  - $L \leq \rho(r)$  everywhere.
  - $L = \rho(r)$  if and only if the radial component of the velocity is zero ( $\dot{r} = 0$ ).

# Reduced phase space



The reduced phase space is under the graph of  $\rho$ .

# Herglotz as convexity

- In the smooth case: Herglotz means convexity of the reduced phase space.

# Herglotz as convexity

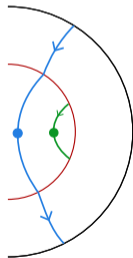
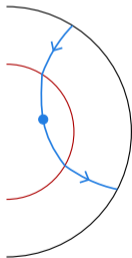
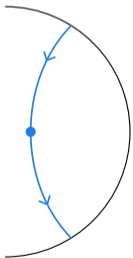
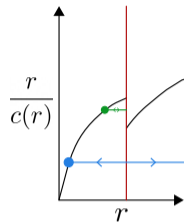
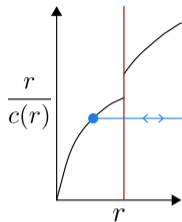
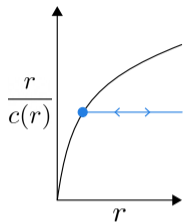
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# Herglotz as convexity

- In the smooth case: Herglotz means convexity of the reduced phase space.
- In the presence of discontinuities: Herglotz in the sense of distributions means convexity.
- For our theorems we only need Herglotz in every layer. Jumps can be in any direction.

# Herglotz as convexity



All of these are fine for us.

# Thank you!

Key ideas:

- Spectrum determines length spectrum through a trace formula.
- Jump discontinuities are not a problem.
- *Should* be implementable.

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